

Pairing and bound states in fermionic systems

Armen Sedrakian

Institute for Theoretical Physics
Tübingen University

Recent Progress in Many Body Theories 14, Barcelona, July 17

Outline

- Alpha particle condensation in dense matter
- Few-body bound states in nuclear medium
- Pairing in nuclear systems
- Weak interactions and neutrino rates

Alpha condensation

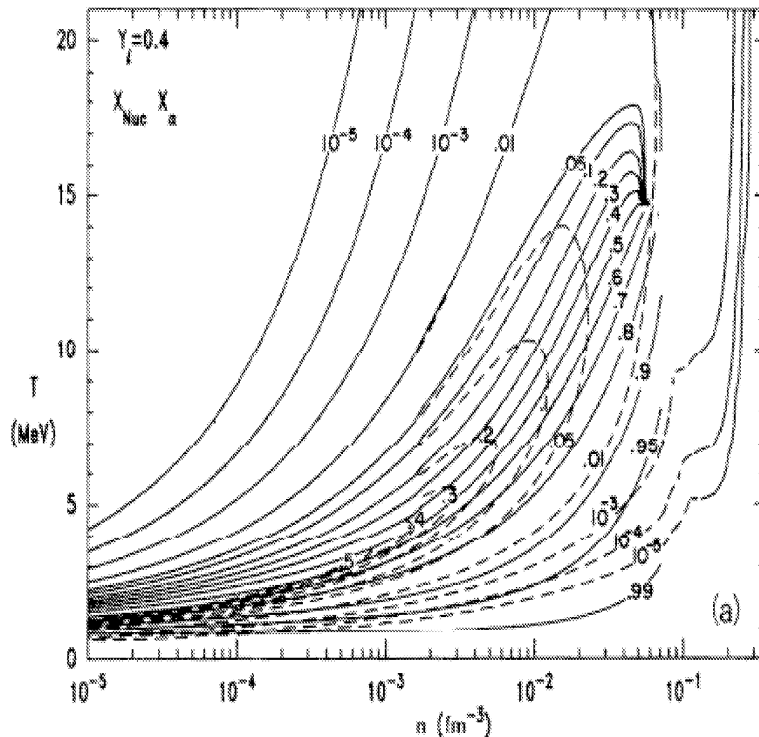
- **Phenomenology of alpha condensation:**
 - The excited states of $4N$ nuclei are well described within the α particle model: elementary degrees of freedom are α 's interacting via a α - α potential: ${}^8\text{Be}$ (unstable) ${}^{12}\text{C}$ (first stable α nucleus), ${}^{16}\text{O}$, ${}^{40}\text{Ca}$.
 - Recent work suggest that these systems are well described by single wave function (BEC in systems with a few particles ?).
 - This motivates the study of Bose-Einstein condensation in infinite alpha matter - start with $N \rightarrow \infty$ system and follow the crossover as N is reduced.

The phase diagram of supernova matter

- Densities $\rho \sim 10^{12} \text{ g cm}^{-3}$; Temperatures $T \leq 10 \text{ MeV}$; content $\rightarrow 15 - 20\% \alpha$ particles

686

J.M. Lattimer et al. / Hot, dense matter



- Quantum degeneracy (BEC)
- Opacities
- Neutrino sphere and signal

From Hamiltonian to effective action

- The non-relativistic theory of bosons can be mapped onto relativistic $O(n)$ model when $T \rightarrow T_c^+$

$$H = \int d^3x \left[\frac{\hbar^2}{2m} \nabla \psi^\dagger(\mathbf{x}) \nabla \psi(\mathbf{x}) - \mu \psi^\dagger(\mathbf{x}) \psi(\mathbf{x}) + \psi^\dagger(\mathbf{x}) \psi(\mathbf{x}) U(\mathbf{x}) \right],$$

$$U(\mathbf{x}) = \int d\mathbf{x}' V_2(\mathbf{x}', \mathbf{x}) \psi^\dagger(\mathbf{x}') \psi(\mathbf{x}') + \int d\mathbf{x}' \int d\mathbf{x}'' V_3(\mathbf{x}, \mathbf{x}', \mathbf{x}'') \psi^\dagger(\mathbf{x}'') \psi(\mathbf{x}') \psi^\dagger(\mathbf{x}'') \psi(\mathbf{x}'),$$

- Expand the fields ψ and $\bar{\psi}$ in Matsubara sums and keep near T_c the zeroth order term (Baym, Blaizot, Zinn-Justin, 1999, Zinn-Justin 1989).
- Introduce new **real fields**: $\psi_{\pm} = \eta(\phi_1 \pm i\phi_2)$.
- The resulting continuum action describes a **classical $O(2)$ symmetric scalar ϕ^6 field theory in 3d**:

$$\mathcal{S}(\phi) = \int d^3x \left\{ \frac{1}{2} \sum_{\nu} [\partial_{\nu} \phi(\mathbf{x})]^2 + \frac{r}{2} \phi(\mathbf{x})^2 + \frac{u}{4!} [\phi(\mathbf{x})^2]^2 + \frac{w}{6!} [\phi(\mathbf{x})^2]^3 \right\}.$$

- Negative quartic, and positive sextic terms!

- Expand the fields ψ and $\bar{\psi}$ in Matsubara sums and keep near T_c the zeroth order term (Baym, Blaizot, Zinn-Justin, 1999, Zinn-Justin 1989).
- Introduce new **real fields**: $\psi_{\pm} = \eta(\phi_1 \pm i\phi_2)$.
- The resulting continuum action describes a **classical $O(2)$ symmetric scalar ϕ^6 field theory in 3d**:

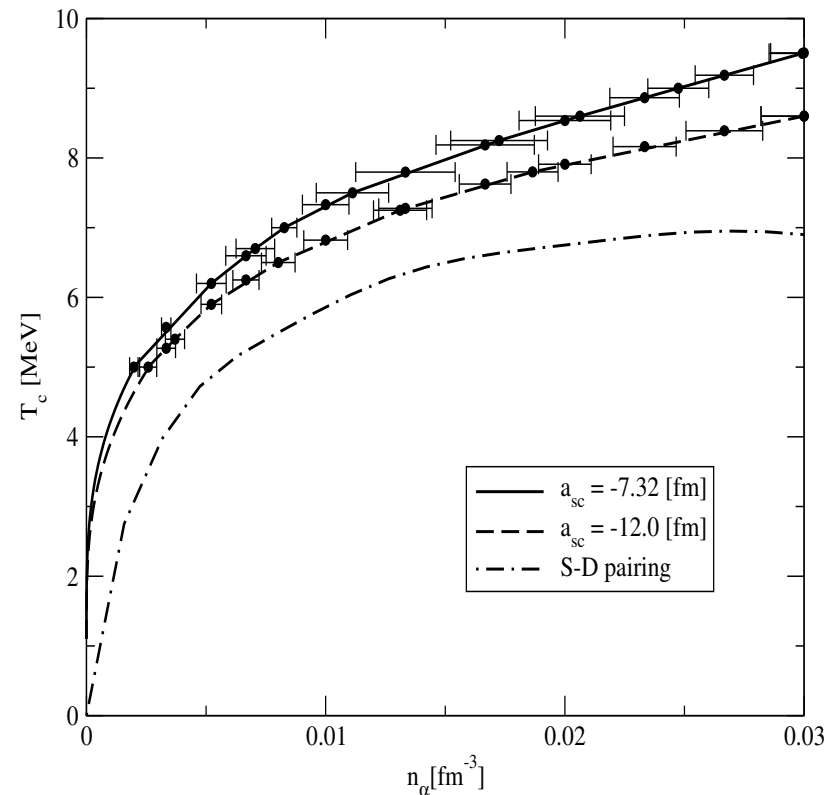
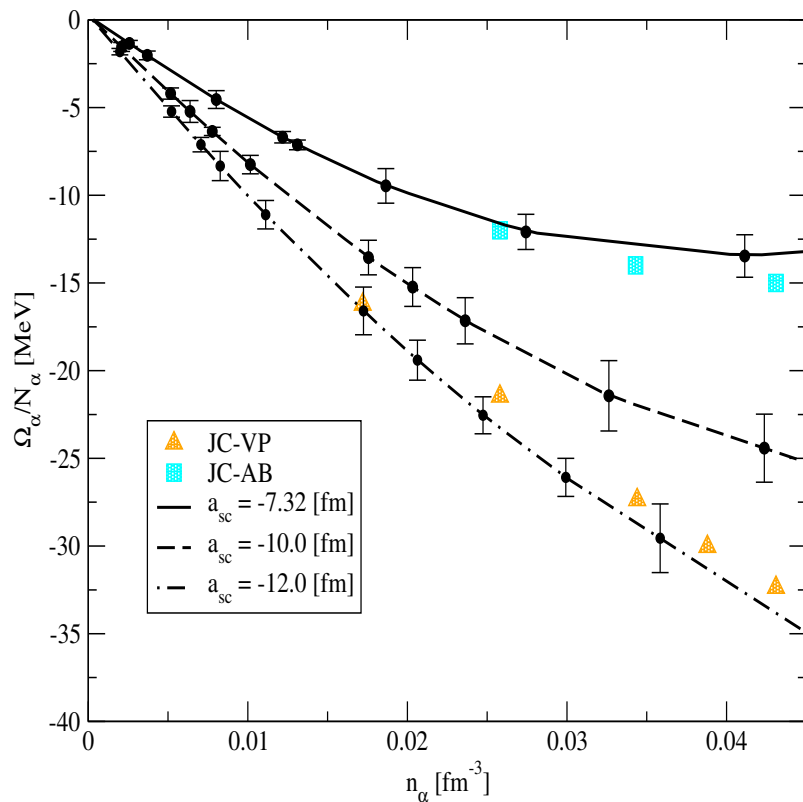
$$\mathcal{S}(\phi) = \int d^3x \left\{ \frac{1}{2} \sum_{\nu} [\partial_{\nu} \phi(\mathbf{x})]^2 + \frac{r}{2} \phi(\mathbf{x})^2 - \frac{u}{4!} [\phi(\mathbf{x})^2]^2 + \frac{w}{6!} [\phi(\mathbf{x})^2]^3 \right\}.$$

- Negative quartic, and positive sextic terms!

Partition function

- Compute the partition function on 3d spatial lattice

$$\mathcal{Z} = \int [d\phi(\mathbf{x})] \exp[-\mathcal{S}(\phi)]$$



- Results could be supplemented by an HNC4/HNC5 calculations at $T = 0$ of condensate fraction.
- Potentially rich astrophysical implications of alpha condensations remain to be studied
- **What about the other light clusters $A = 2, 3$?**
 - The deuteron and triton are less bound (-2.2 MeV) and (-8.4 MeV) and thus less stable.
 - Weakly bound states are ideal sources of opacity!
example: H^- opacity is dominant in ordinary stars.

The three-body problem in background medium

- Work within the Kadanoff-Baym real-time GF method → two and three-point functions are defined on the Schwinger-Keldysh contour
- The three-body equation for the \mathcal{T} -matrix

$$\mathcal{T} = \mathcal{V} + \mathcal{V} \mathcal{G} \mathcal{V} = \mathcal{V} + \mathcal{V} \mathcal{G}_0 \mathcal{T},$$

where the interaction $\mathcal{V} = V_{12} + V_{23} + V_{13}$

- Reformulate the problem: $\mathcal{T} = \mathcal{T}^{(1)} + \mathcal{T}^{(2)} + \mathcal{T}^{(3)}$

$$\mathcal{T}^{(k)} = \mathcal{V}_{ij} + \mathcal{V}_{ij} \mathcal{G}_0 \mathcal{T} \quad ijk = 123, 231, 312.$$

Define: $\mathcal{T}_{ij} = \mathcal{V}_{ij} + \mathcal{V}_{ij} \mathcal{G}_0 \mathcal{T}_{ij}$ and eliminate the potentials

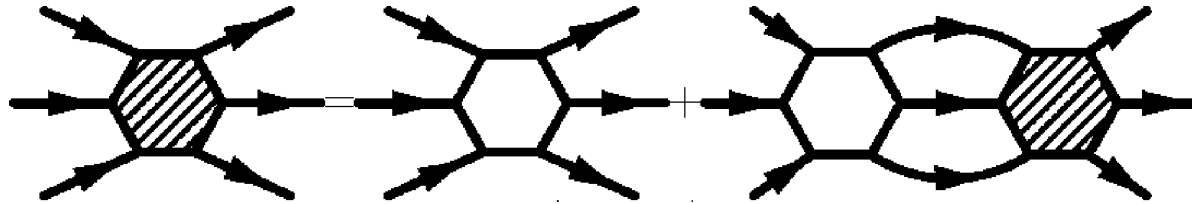
- Non-singular three-body equations (Bethe-Faddeev) with the time structure of the three-body T -matrix

$$\mathcal{T}^{R(1)}(t, t') = \mathcal{T}_{23}^R(t, t') + \int [\mathcal{T}^{R(2)}(t, \bar{t}) + \mathcal{T}^{R(3)}(t, \bar{t})] \mathcal{G}_0^R(\bar{t}, t'') \mathcal{T}_{23}^R(t'', t') d\bar{t} dt'',$$

- Many particle-hole channels (new aspect of medium physics)

$$\mathcal{G}_0^R(t_1, t_2) = \theta(t_1 - t_2) \left\{ \begin{array}{ll} G^> G^> G^>(t_1, t_2) - (> \leftrightarrow <) & (3p) \\ G^> G^> G^<(t_1, t_2) - (> \leftrightarrow <) & (2ph) \\ G^> G^< G^<(t_1, t_2) - (> \leftrightarrow <) & (p2h) \\ G^< G^< G^<(t_1, t_2) - (> \leftrightarrow <) & (3h) \end{array} \right.$$

Particle-hole content of the T -matrix



- 3-particle – 3-hole scattering T -matrix

$$\mathcal{T}^{R(1)} = \mathcal{T}_{23}^R + \int \left[\mathcal{T}^{R(2)} + \mathcal{T}^{R(3)} \right] \frac{Q_3(\Omega')}{\Omega - \Omega' + i\eta} \mathcal{T}_{23}^R(\Omega') d\Omega',$$

- 3-body Pauli-blocking: $\bar{f}_F = 1 - f_F$

$$Q_3(p_\alpha, p_\beta, p_\gamma) = \bar{f}_F(p_\alpha) \bar{f}_F(p_\beta) \bar{f}_F(p_\gamma) - f_F(p_\alpha) f_F(p_\beta) f_F(p_\gamma).$$

p_α are spanned in terms of Jacobi coordinates.

Bound states in background medium

- Bound state wave-function

$$\Psi = \psi^{(1)} + \psi^{(2)} + \psi^{(3)}; \quad \psi^{(k)} = \mathcal{G}_0 T_{ij} (\psi^{(i)} + \psi^{(j)}). \quad (-8)$$

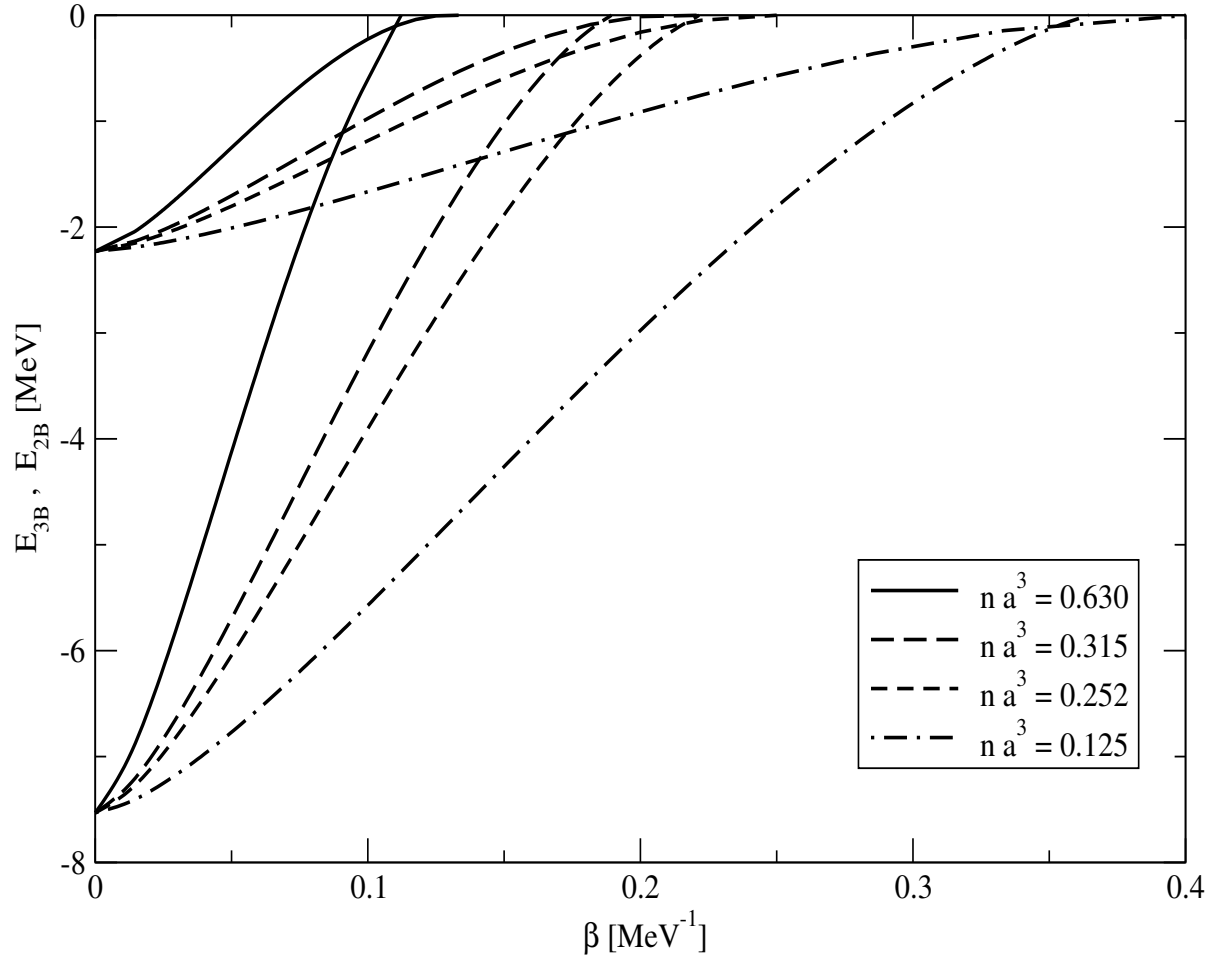
- Need the channel T -matrix

$$\begin{aligned} T^R(\vec{p}, \vec{p}'; \vec{P}, E) \\ = V(\vec{p}, \vec{p}') + \int \frac{d\vec{p}''}{(2\pi)^3} V(\vec{p}, \vec{p}'') G_0^R(\vec{p}'', \vec{P}, E) T^R(\vec{p}'', \vec{p}'; \vec{P}, E) \end{aligned}$$

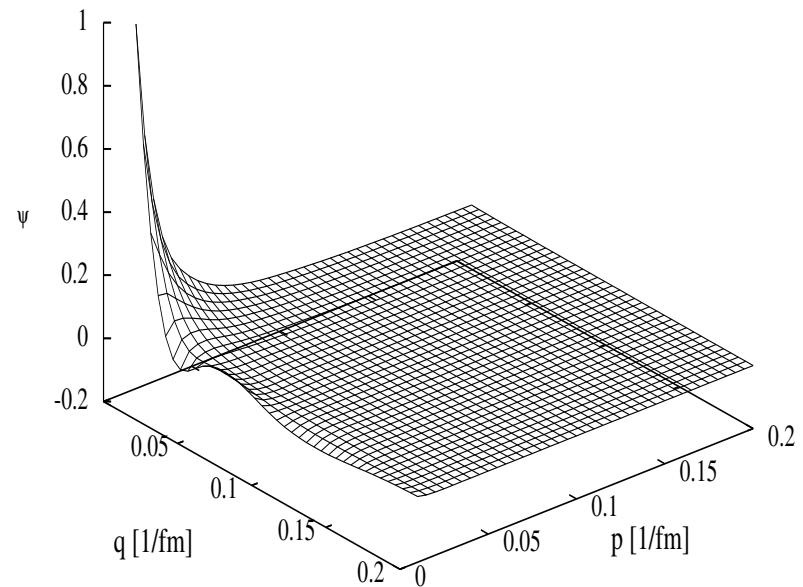
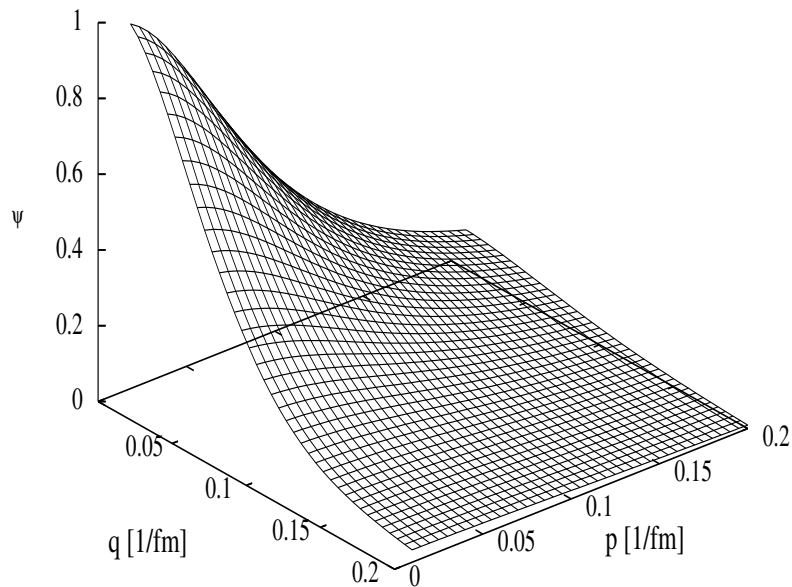
$$G_0^R(\vec{k}_1, \vec{k}_2, E) = \frac{Q_2(\vec{k}_1, \vec{k}_2)}{E - \epsilon(\vec{k}_1) - \epsilon(\vec{k}_2) + i\eta}, \quad (-10)$$

Poles $\rightarrow A = 2$ binding energies $E_B(T, \rho)$.

Background dependent binding energies of triton in nuclear matter



The ratio $\eta = E_{3B}(T)/E_{2B}(T)$ is independent of temperature.



Squeezing the wave function in the momentum space: signal of a quantum phase transition to the unbound state.

What happens beyond the extinction point?

- The deuteron crosses over to become a Cooper pair
- Precritical three-body correlations for $T > T_c$?

Nozieres-Schmitt-Rink solutions

The evolution deuteron \rightarrow Cooper pair is well described by the mean-field BCS theory (coupled ${}^3S_1 - {}^3D_1$ channels)

$$\Delta^{(i)}(p, \Lambda) = \int^{\Lambda} \frac{dp' p'^2}{(2\pi)^3} V^{3SD1}(p, p') \times \frac{\Delta^{(i-1)}(p', \Lambda)}{\sqrt{E_p^2 + D^{(i-1)}(p', \Lambda)^2}} [f(\omega_+) - f(\omega_-)]. \quad (-11)$$

$$n = 4 \int \frac{d^3p}{(2\pi)^3} [u_p^2 f(\omega_+) + v_p^2 f(\omega_-)]. \quad (-11)$$

Solved iteratively with a running cut-off Λ ; the iteration is stopped when $d\Delta/d\Lambda = 0$.

BCS-BEC crossover

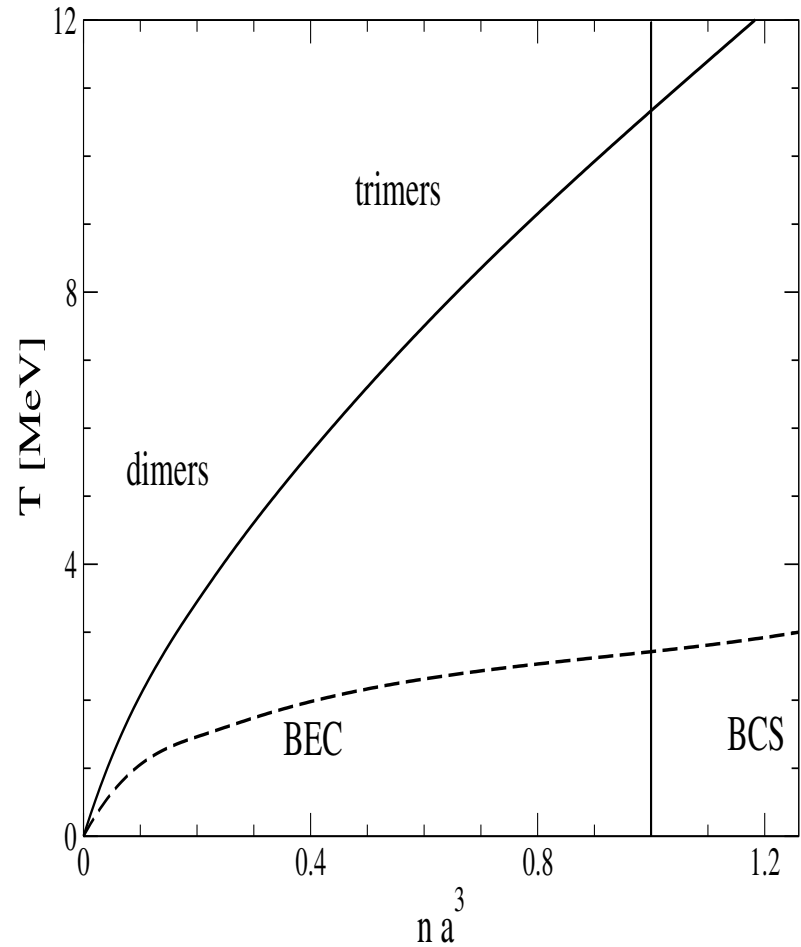
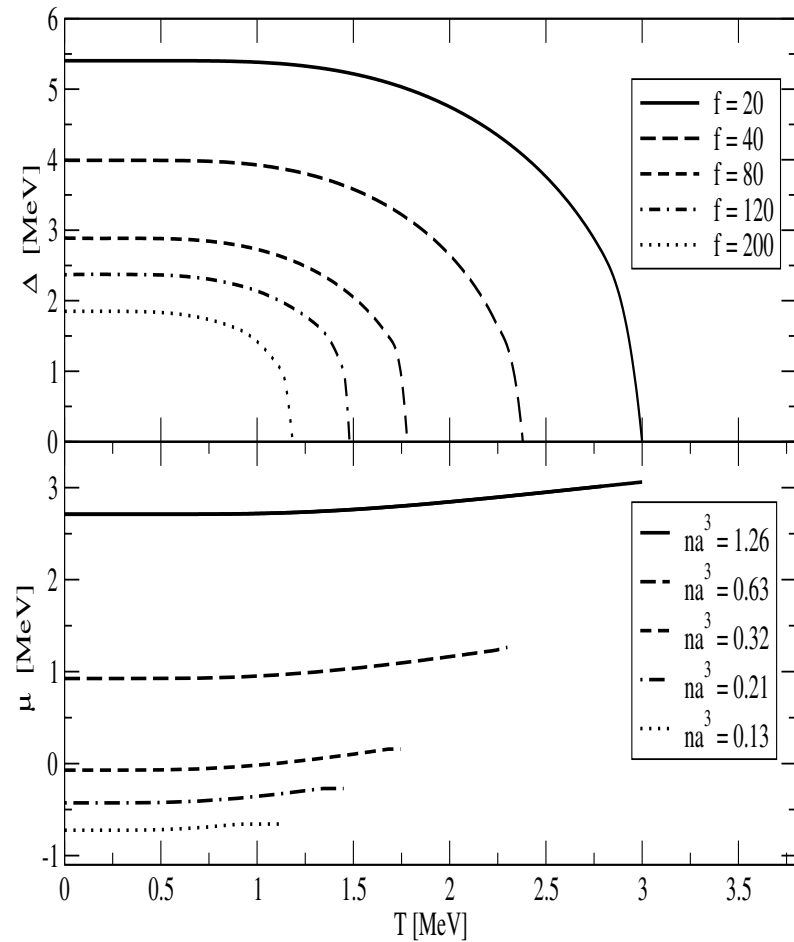
- Nozières-Schmitt-Rink conjecture: the BCS theory smoothly interpolates between the weak and strong couplings
- In the BEC limit the pair-wave function goes over to the Schrödinger equation

$$\psi(k) = \langle a_{n,\vec{k}}^\dagger a_{p,-\vec{k}}^\dagger \rangle = \frac{\Delta(k)}{2E_k} [1 - f(E_k^+) - f(E_k^-)] ,$$

$$\frac{k^2}{m} \psi(k) + [1 - f(E_k^+) - f(E_k^-)] \sum_{k'} V(k, k') \psi_{l'}(k') = 2\mu \psi(k)$$

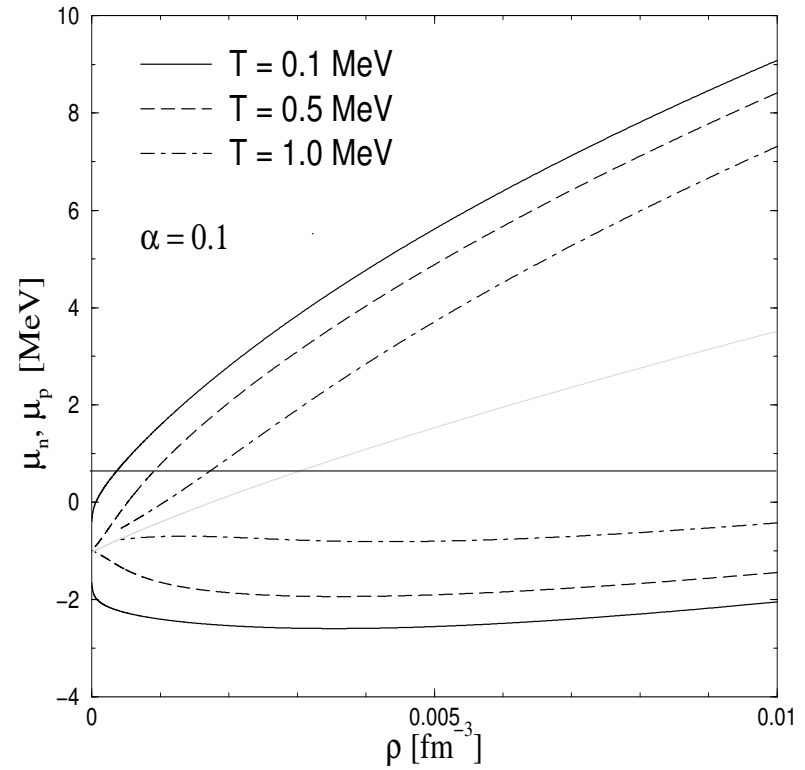
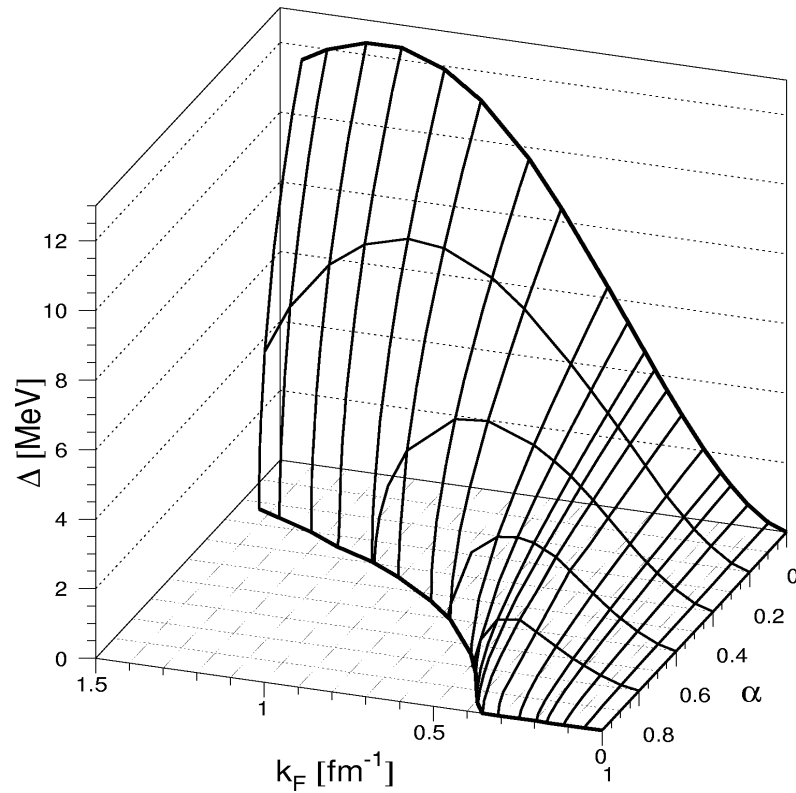
- In unbalanced systems phases with broken space symmetries intervene.

BCS-BEC crossover



At the cross-over μ sign changes, coherence length $\xi \sim n^{1/3}$.

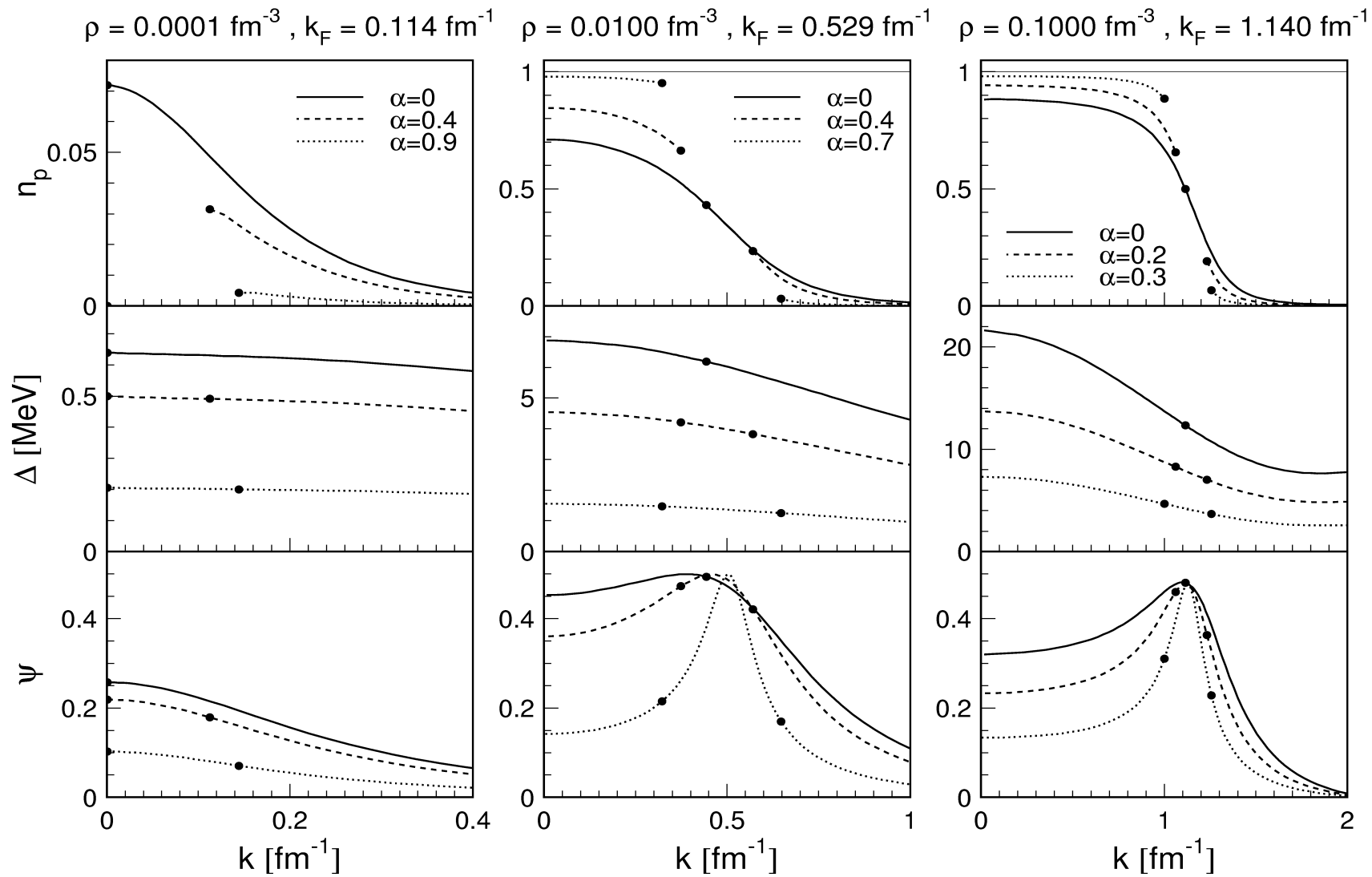
BCS-BEC crossover in asymmetric systems



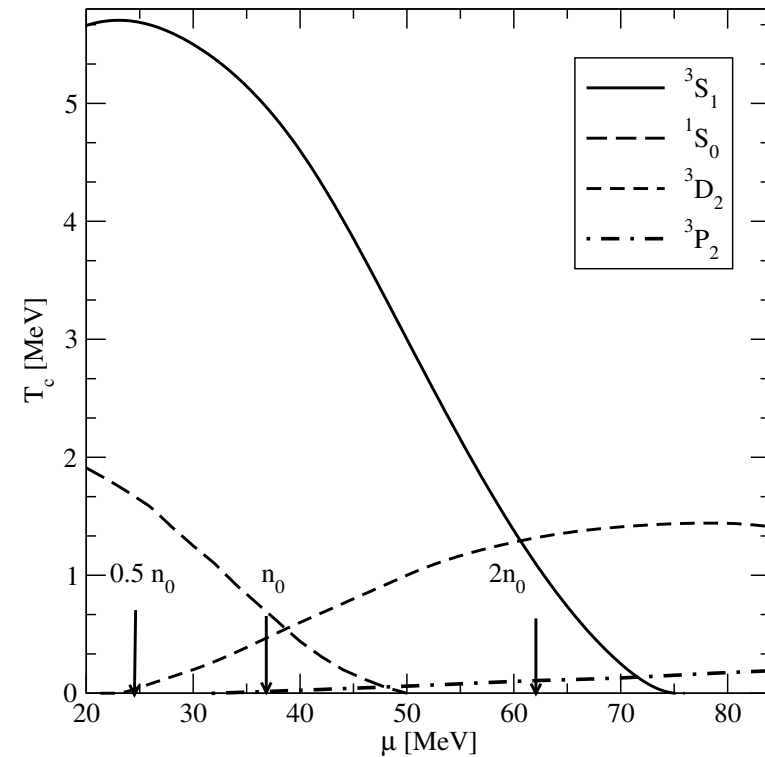
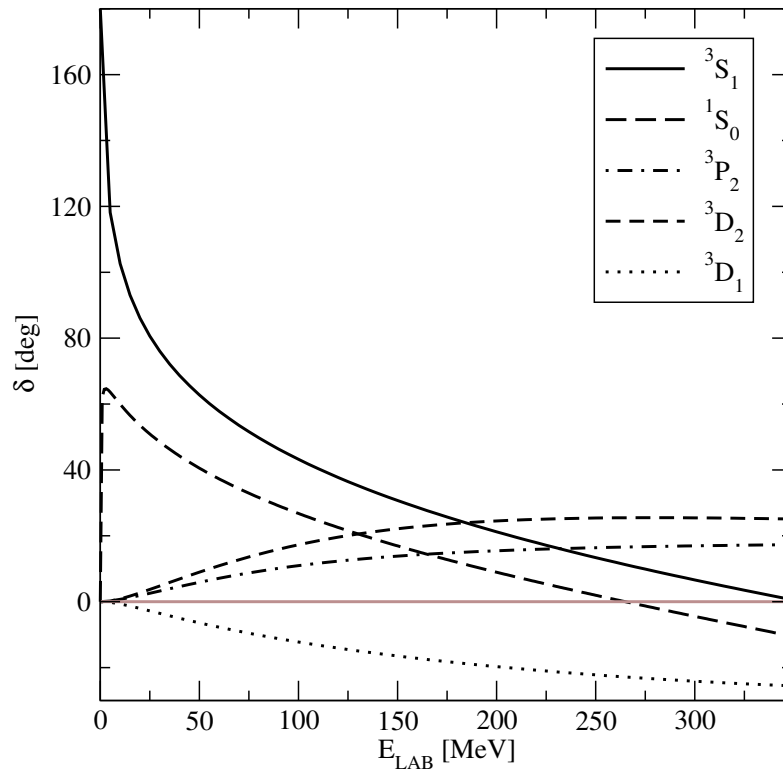
U. Lombardo, P. Nozieres, P. Schuck, H.-J. Schulze, A. Sedrakian, Phys.Rev. C64 (2001)

064314

Crossover continued



Nuclear systems



Left panel. Dependence of the experimental scattering phase shifts in the 3S_1 , 3P_2 , 3D_2 , and 3D_1 partial waves on the laboratory energy. *Right panel.* The dependence of the critical temperatures of superfluid phase transitions in the attractive channels on the chemical potential.

- Pairing in neutron matter with retardation - strong coupling Eliashberg theory
 - time-non-local interactions - gaps are then energy dependent
 - keep same approximations for normal and anomalous self-energies (Ward identities)
 - possible in medium modifications of the meson dynamics (e.g. precursor of pion condensation).
- Long + short range components:

$$\begin{aligned}
 H_{\pi NN} &= -\frac{f_{\pi}}{m_{\pi}}(\boldsymbol{\sigma} \cdot \boldsymbol{\nabla})(\boldsymbol{\tau} \cdot \boldsymbol{\phi}) \\
 &+ \left[g' \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2 + h' \boldsymbol{\sigma}_1 \cdot \boldsymbol{\sigma}_2 S_{12}(\boldsymbol{n}) \right] \boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_2,
 \end{aligned}$$

DSE in nucleonic sector

- Normal propagators:

$$G^R(p) = \frac{\omega Z(p) + \xi_p^*}{(\omega + i\eta)^2 Z(p)^2 - \xi_p^{*2} - \Delta^R(p)^2},$$

- Anomalous propagators:

$$F^R(p) = -\frac{\Delta^R(p)}{(\omega + i\eta)^2 Z(p)^2 - \xi_p^{*2} - \Delta^R(p)^2},$$

- Normal self-energies: $\Sigma^R = \Sigma_S^R + \Sigma_A^R$

$$\xi^*(p) = \xi(p) + \Sigma_S^R(p)$$

- Wave-function renormalization: $Z(p) = 1 - \omega^{-1} \Sigma_A^R(\omega)$.

DSE in bosonic sector

- **DSE for mesons:** $\hat{D}(p) = \hat{D}_0(p) + \hat{D}_0(p)\hat{\Pi}(p)\hat{D}(p)$.
- **Polarization tensor and the vertex:**

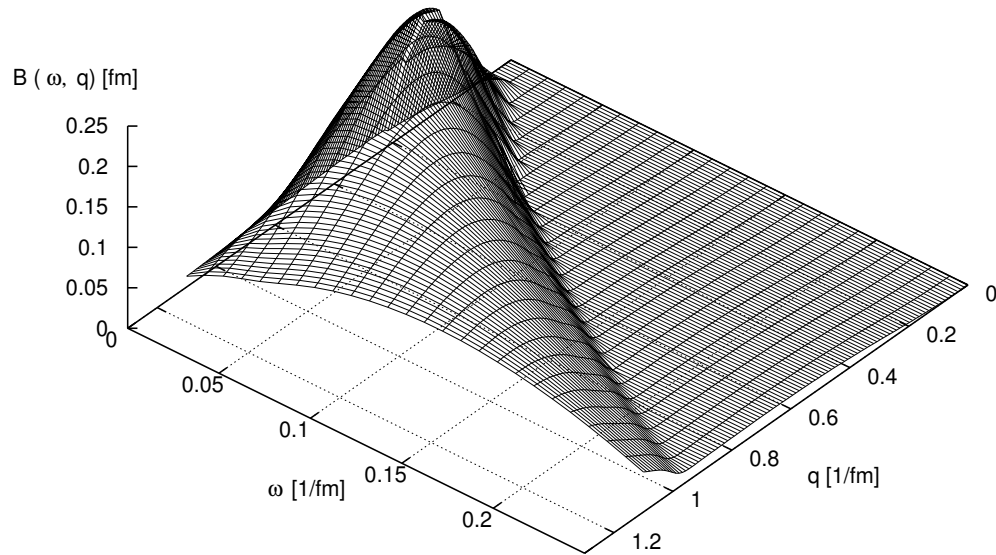
$$\hat{\Pi}(q) = -\text{Tr} \int \frac{d^4 p}{(2\pi)^4} i \hat{\Gamma}_0(q) \hat{G}(p+q) \hat{G}(-p) \hat{\Gamma}(q),$$

$$\hat{\Gamma}(q) = \hat{\Gamma}_0(q) + \text{Tr} \int \frac{d^4 p}{(2\pi)^4} i \hat{\Gamma}_1(q) \hat{G}(p+q) \hat{G}(-p) \hat{\Gamma}(q).$$

- **Pion spectral function:**

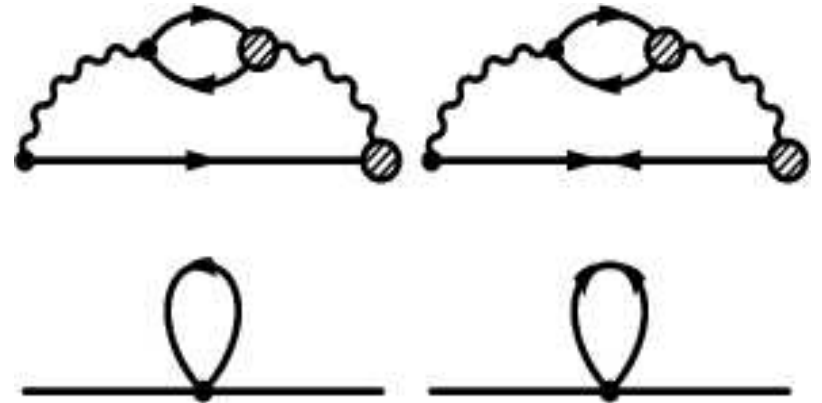
$$B(q) = \frac{-2\text{Im}\Pi^R(q)}{[\omega^2 - \vec{q}^2 - m_\pi^2 - \text{Re}\Pi^R(q)]^2 + [\text{Im}\Pi^R(q)]^2}.$$

Pion spectral function



← Pion spectral function in cold neutron matter

Hartree and Fock contributions to the self-energies of neutrons →



Retarded (Fock) part of the self-energies

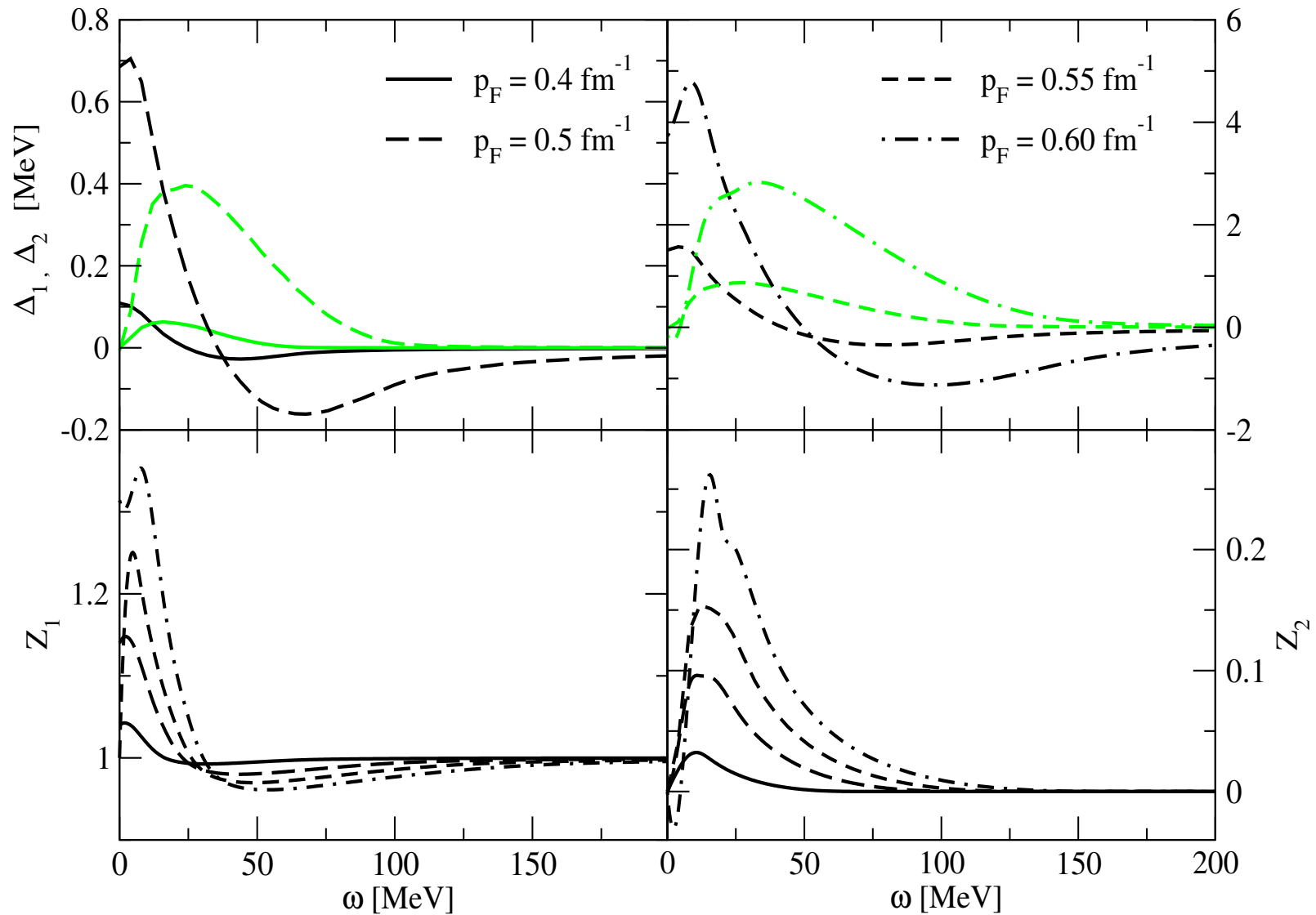
$$\Sigma^R(\omega, \vec{p}) = \text{Tr} \int \frac{d^3 q d\varepsilon}{(2\pi)^4} \Gamma_0(\vec{q}) A_G(\varepsilon, \vec{p} - \vec{q}) C(\omega, \varepsilon, \vec{q}) \Gamma(\vec{q}),$$

$$\Delta^R(\omega, \vec{p}) = \text{Tr} \int \frac{d^3 q d\varepsilon}{(2\pi)^4} \Gamma_0(\vec{q}) A_F(\varepsilon, \vec{p} - \vec{q}) C(\omega, \varepsilon, \vec{q}) \Gamma(\vec{q}),$$

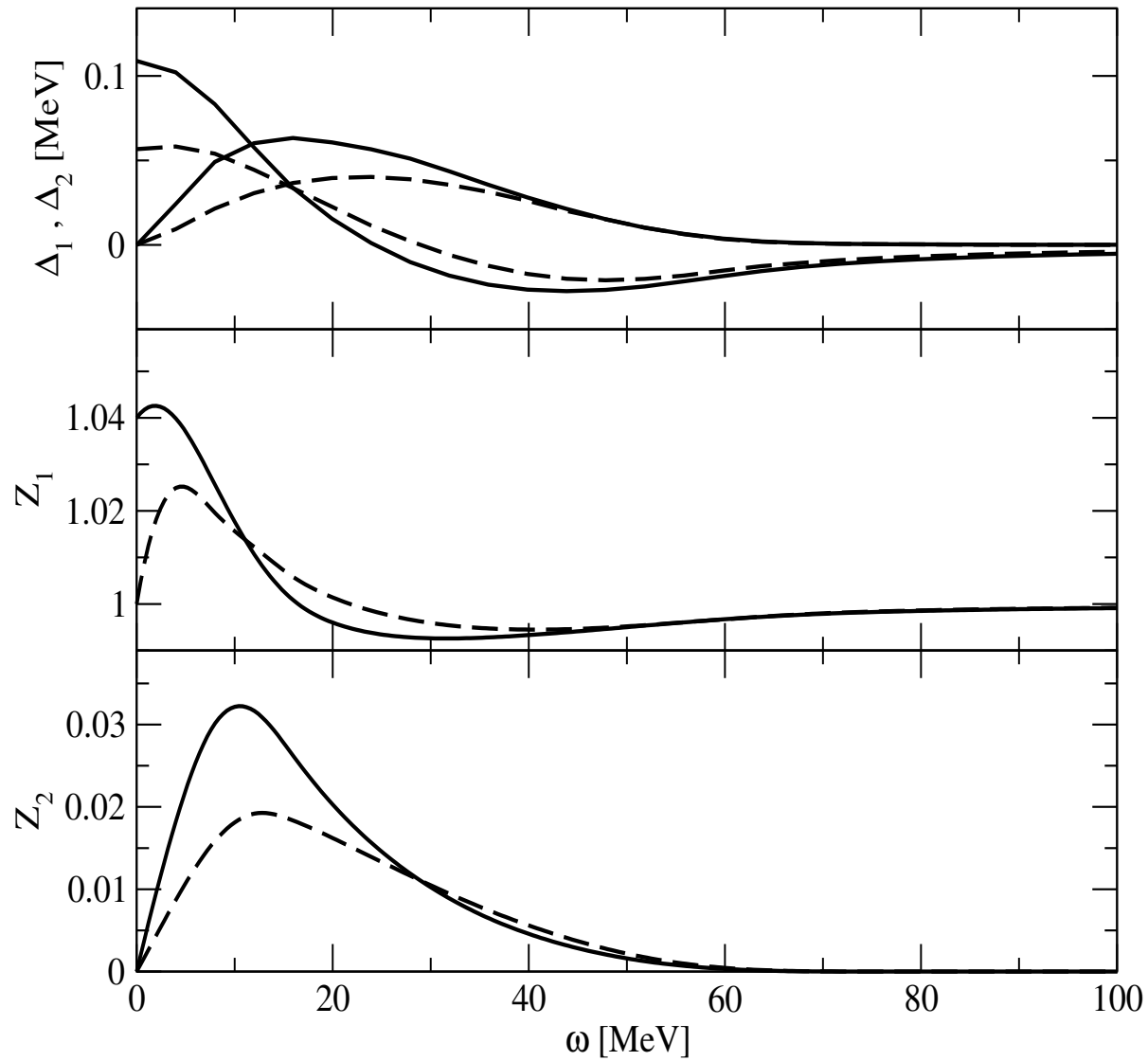
$$C = \int_0^\infty \frac{d\omega'}{2\pi} B(\omega', \vec{q}) \left[\frac{f(\varepsilon) + g(\omega')}{\varepsilon - \omega' - \omega - i\eta} + \frac{1 - f(\varepsilon) + g(\omega')}{\varepsilon + \omega' - \omega - i\eta} \right]$$

- Need to solve self-consistently 4 integral equations (= 2 complex equation for the normal and anomalous self-energies).
- The kernels of these integral equations are singular either at the boundary or within the integration range.
- Iterative procedure; good convergence after 15-20 iterations.

Off-shell characteristics: gap and wave-function renormalization



Different approximations to the spectral function



Effects of pairing on neutrino radiation from NS

Processes on fermions

- Neutral current processes (Z_0 exchange)

$$\begin{cases} f_1 \rightarrow f_2 + \nu_f + \bar{\nu}_f & (\text{brems}) \\ f_1 + f'_1 \rightarrow f_2 + f'_2 + \nu_f + \bar{\nu}_f \end{cases} \quad (-21)$$

- Charged current processes (W^\pm exchange)

$$\begin{cases} f_1 \rightarrow f_2 + e + \bar{\nu}_e & (\text{Urca}) \\ f_1 + f'_1 \rightarrow f_2 + f'_2 + e + \bar{\nu}_e \end{cases} \quad (-22)$$

Transport equations

- ν and $\bar{\nu}$ - Boltzmann equations with KB collision integrals

$$\begin{aligned} & \left[\partial_t + \vec{\partial}_q \omega_\nu(\vec{q}) \vec{\partial}_x \right] f_\nu(\vec{q}, x) \\ &= \int_0^\infty \frac{dq_0}{2\pi} \text{Tr} \left[\Omega^<(q, x) S_0^>(q, x) - \Omega^>(q, x) S_0^<(q, x) \right], \end{aligned}$$

- ν -quasiparticle propagators:

$$\begin{aligned} S_0^<(q, x) = \frac{i\pi \not{q}}{\omega_\nu(\vec{q})} & \left[\delta(q_0 - \omega_\nu(\vec{q})) f_\nu(q, x) \right. \\ & \left. - \delta(q_0 + \omega_\nu(\vec{q})) (1 - f_{\bar{\nu}}(-q, x)) \right]. \end{aligned}$$

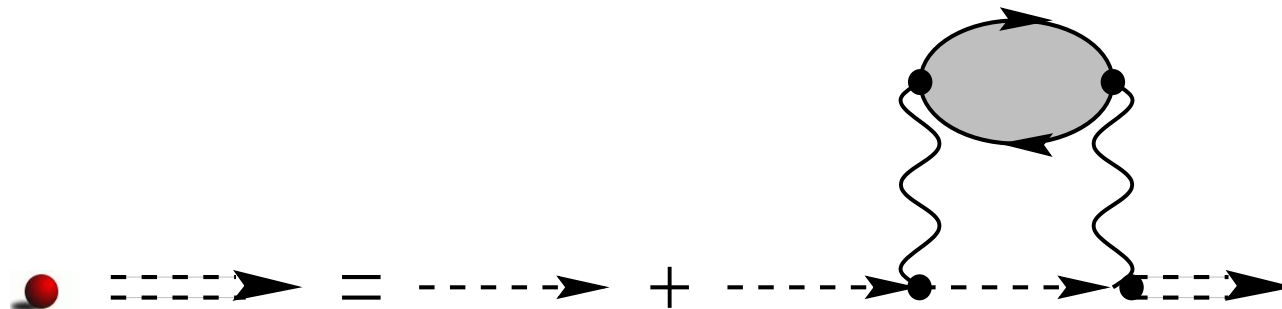
- definition of the Poisson bracket

$$\{f, g\}_{P.B.} = \partial_\omega f \partial_t g - \partial_t f \partial_\omega g - \partial_{\vec{p}} f \partial_{\vec{r}} g + \partial_{\vec{r}} f \partial_{\vec{p}} g.$$

Self-energies

- ν and $\bar{\nu}$ -self-energies (second order in weak force)

$$-i\Omega^{>,<}(q_1, x) = \int \frac{d^4 q}{(2\pi)^4} \frac{d^4 q_2}{(2\pi)^4} (2\pi)^4 \delta^4(q_1 - q_2 - q) \\ i\Gamma_{Lq}^\mu iS_0^<(q_2, x) i\Gamma_{Lq}^{\dagger\lambda} i\Pi_{\mu\lambda}^{>,<}(q, x), \quad (-26)$$



- the central problem is to compute the polarization tensor!

Bremsstrahlung emissivity

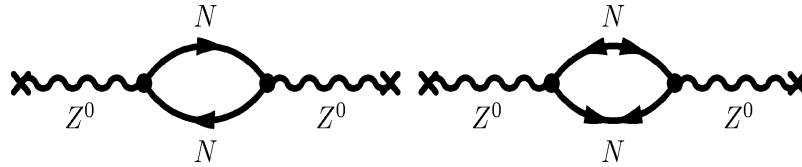
- energy loss per unit time and volume

$$\epsilon_{\nu\bar{\nu}} = \frac{d}{dt} \int \frac{d^3q}{(2\pi)^3} [f_\nu(\vec{q}) + f_{\bar{\nu}}(\vec{q})] \omega_\nu(\vec{q}) \quad (-27)$$

- expressed through the collision integrals

$$\begin{aligned} \epsilon_{\nu\bar{\nu}} = & -2 \left(\frac{G}{2\sqrt{2}} \right)^2 \sum_f \int \frac{d^3q_2}{(2\pi)^3 2\omega_\nu(\vec{q}_2)} \int \frac{d^3q_1}{(2\pi)^3 2\omega_\nu(\vec{q}_1)} \int \frac{d^4q}{(2\pi)^4} \\ & (2\pi)^4 \delta^3(\vec{q}_1 + \vec{q}_2 - \vec{q}) \delta(\omega_\nu(\vec{q}_1) + \omega_\nu(\vec{q}_2) - q_0) [\omega_\nu(\vec{q}_1) + \omega_\nu(\vec{q}_2)] \\ & g_B(q_0) [1 - f_\nu(\omega_\nu(\vec{q}_1))] [1 - f_{\bar{\nu}}(\omega_\nu(\vec{q}_2))] \Lambda^{\mu\lambda}(q_1, q_2) \Im \Pi_{\mu\lambda}^R(q). \end{aligned}$$

Neutral current pair-breaking processes



- The one-loop contribution to the polarization tensor in the superfluid matter

$$\epsilon_{\nu\bar{\nu}} = \frac{G^2 c_V^2}{240\pi^3} \nu(p_F) T^7 I(\zeta) \equiv \epsilon_0 I(\zeta),$$

$$I(\zeta) = \zeta^7 \int_0^\infty d\phi (\cosh \phi)^5 f(\zeta \cosh \phi)^2, \quad \zeta = 2\Delta(T)/T$$

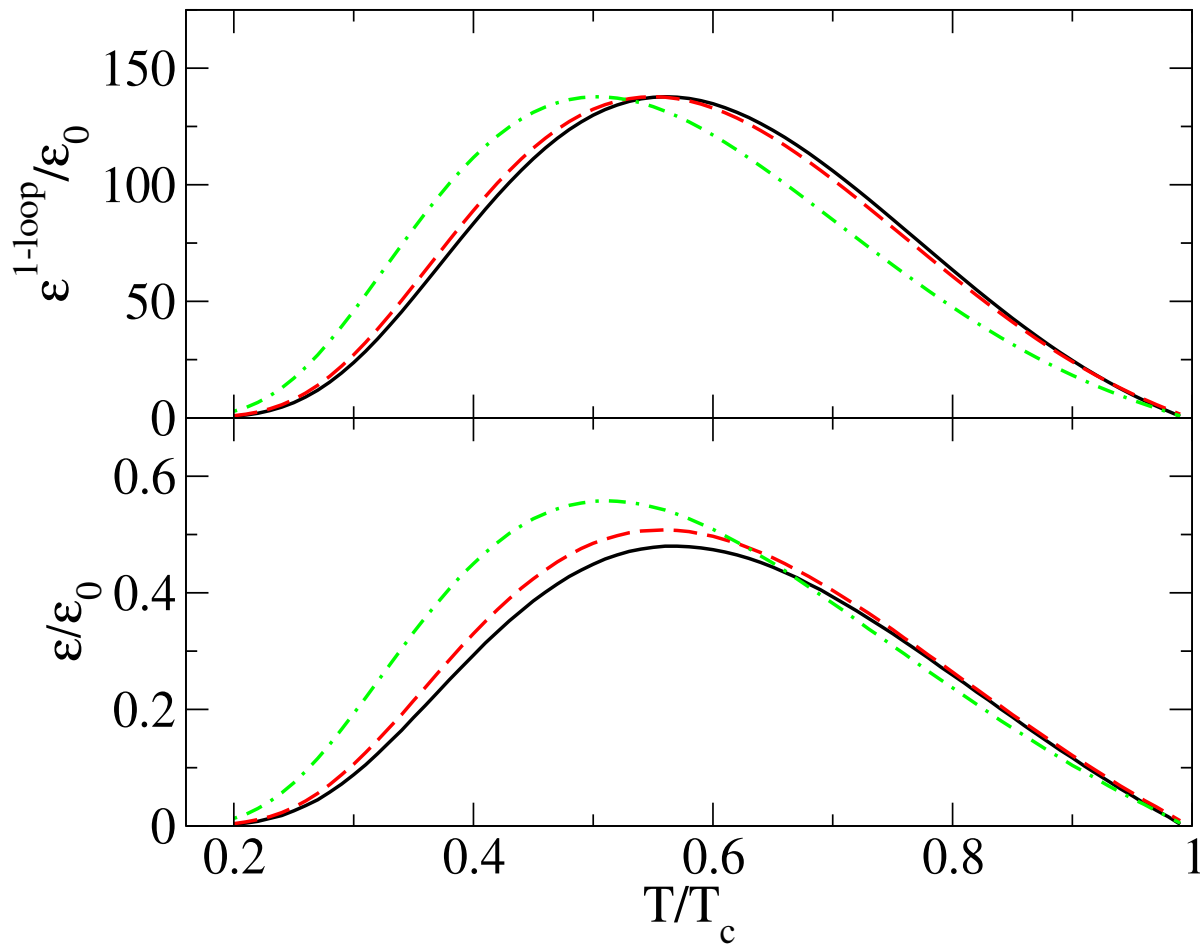
Need to go beyond the one loop - RPA

- The vertex functions in the superfluid phase

$$\begin{aligned}
 \Gamma_1 &= \text{[triangle with solid black fill]} + \text{[circle with dashed line]} + \text{[circle with solid line]} + \text{[circle with solid line]} + \text{[circle with solid line]} \\
 \Gamma_2 &= \text{[circle with dashed line]} + \text{[circle with solid line]} + \text{[circle with solid line]} + \text{[circle with solid line]} \\
 \Gamma_3 &= \text{[circle with dashed line]} + \text{[circle with solid line]} + \text{[circle with solid line]} + \text{[circle with solid line]}
 \end{aligned}$$

- Contributions to the polarization tensor

$$\Pi^{(a)} + \Pi^{(b)} + \Pi^{(c)} + \Pi^{(d)}$$



f -sum rule and gauge invariance

$$\int_{-\infty}^{\infty} d\omega \operatorname{Im} \Pi^R(\mathbf{q}, \omega) = \frac{q^2}{2m}; \quad q^\mu \Pi_{\mu\nu}^R(\mathbf{q}, \omega) = 0$$

Thanks to the people who contributed to the research reported here

John Clark (St. Louis)

Herbert Müther (Tübingen)

Peter Schuck (Orsay)

Umberto Lombardo and Hans Schulze (Catania)

Philippe Nozieres (Grenoble)

Alex Dieperink (Groningen)

